A Preliminary Version of a Catalog of High-Low-Mass X-ray Binaries

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Abstract—We introduce a catalogue of X-ray binaries. The full catalogue is available at http://astroa.physics.metu.edu.tr/XRBC/. The aim of this catalogue is to provide basic information about X-ray sources and their counterparts. The catalog contains positions, information about distances and counterparts, spectral, photometric and timing properties of the X-ray sources in the X-ray and optical bands, as well as references. The list of references, comments and the catalog as a whole will be periodically updated to include the most up-to-date information about all X-ray binaries and guide users to the recent literature on individual sources. In some cases, there is some doubt about the nature of an X-ray source, which has been noted. The sources are ordered according to right ascension. Here, we present a guide to the organization of the full catalogue. The full catalogue is represented here with an excerpt containing two HMXBs and two LMXBs as examples. © 2000 MAIK "Nauka/Interperiodica".

Key words: catalogs, X-ray binaries, neutron stars, X-rays

INTRODUCTION

The discovery of the first point X-ray source Sco X-1 (Giacconi et al. 1963) outside the solar system was a surprise, because the only previously known stellar X-ray emission, that of the solar corona, did not lead to an expectation of bright point X-ray sources. The general properties of point X-ray sources in the Galaxy were understood in terms of accretion onto compact objects (Zel'dovich and Guseinov 1965; Shklovskii 1967; Guseinov 1970) just before the first UHURU satellite sky survey in 1971. The results of the UHURU sky survey (Forman et al. 1978) were followed by publications from various rocket and satellite observations (Ariel V sky survey, McHardy et al. 1981; Warwick et al. 1981; MIT OSO-7, Markert et al. 1979; HEAO A-1 sky survey, Wood et al. 1984), and catalogs were compiled to bring together data from different experiments (Amnuel et al. 1979, 1982).

The first X-ray transient (nova) Cen X-2 was discovered from a rocket (Harries 1967). The first X-ray pulsar Cen X-3 was observed in 1971 (Giacconi *et al.* 1971; Schreier 1972). The first burster LMXB 1820-303 was discovered in 1976 (Grindlay *et al.* 1976); Cyg X-1 was discovered in 1965 and discussed as a black-hole candidate in 1972. The first quasi-periodic oscillation (QPO) behavior was also noted from Cyg X-1 (Frontera and Fuligni 1975). The classic QPO phenomenon from LMXBs was first discovered by Van der Klis *et al.* (1985) from the source LMXB 1758-250.

The first comprehensive catalog of X-ray binaries was presented by Van Paradijs (1995). Our catalog includes, as X-ray binaries, the point X-ray sources powered by accretion onto neutron stars and black holes. The weak sources powered by stellar coronas, Wolf-Rayet stars, and cataclysmic variables are not included. Among those X-ray binaries presented in the catalog, there are 78 HMXBs (35 and 5 of which are pulsars and possible pulsars, respectively) and 138 LMXBs (only 9 of which are pulsars), 89 transient and possible transient sources, and 25 possible black holes (6 HMXBs+19 LMXBs). There are also 43 bursters (all of which are Galactic LMXBs) and 21 QPOs. The present catalog extends the work by Van Paradijs, tabulates further information on the compact objects and their companions, as discussed below, as well as attempts to provide a more comprehensive guide to the literature. Where different authors obtained different values for system parameters, we have quoted all parameter values together with the references to guide the reader to the relevant literature for the derivation of the parameters. We introduce the catalog with four examples given in tabular form below.

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Data for 2 HMXBs and 2 LMX

			Table 1 (samp	ple)		
Name	Location	Туре	RA (2000)	Dec (2000)	RA (1950)	Dec (1950)
0115+634		Р, Т	01 ⁿ 18 ^m 32 ^s	63°44′24″	01 ⁿ 15 ^m 13 ^s .8	63°28′38″
0115–737	SMC	Р	01 17 05	-73 2635	01 15 45.6	-73 42 22
J0422+32		Q, R	04 21 42.7	32 54 27.1	04 18 29.9	32 47 24
0620-003		R	06 22 44.4	00 20 44.7	06 20 11.1	-00 19 11

			Tab	le 2 (sample)				
Name	l ^{II}	b^{II}	$P_{\rm orb}$, days	P _{pulse} , s	P _{rem}	<i>kT</i> , keV	α	Distance, kpc
0115+634	125.9	1.0	24.309	3.61451	*			*
0115–737	300.4	-43.6	3.89239	0.717	*	0.18*		57
J0422+32	165.9	-11.9	0.212140			2		1.2*
0620-003	210	-6.5	0.325		*	*	~1.6	1*

			Table 3 ((sample)			
Name	$\log N_{\rm H}, {\rm cm}^{-2}$	A_V , mag.	$L_{\rm X}$, erg s ⁻¹	$L_{\rm X}/L_0$	SpecT	V, mag.	<i>B–V</i> , mag.
0115+634	22.08	>5	*	1.7	OBe	14.8*	1.61*
0115-737	<21.60	0.1	6.0×10^{38}	*	B0 Ib	13.3	-0.14
J0422+32	<21.7*	1.2*	2.7×10^{37}	*	M2 V	22.24*	0.15*
0620–003	21.2*	1.2	$1.0 \times 10^{38*}$	200	K4–K7 V	18.3	1.4

Table 4 (sample)
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Name	<i>U–B</i> , mag.	<i>E</i> (<i>B</i> – <i>V</i>), mag.	$M_{\rm X}, M_{\odot}$	$M_{ m opt}, M_{\odot}$	$f_{\rm X}(M), M_{\odot}$	$f_{\rm opt}(M), M_{\odot}$	e _{opt}
0115+634	0.3	1.7		>20*	5		0.34
0115-737	-0.98	~0.1*	1.17*	15.2*	10.8		<10 ⁻⁵ *
J0422+32	-0.5	0.2*	>3*	0.39*	3.1*	0.9	
0620-003	-0.8	0.35*	5.17	*		2.9	

			Table 5 (sample)			
Name	$K_{\rm X}$, km s ⁻¹	$K_{\rm opt}$, km s ⁻¹	γ, km s ⁻¹	$a_1 \sin i$, light-s	$a_2 \sin i$, light-s	i, deg
0115+634	133.7		365	140		
0115-737	299.5	*	180	53.46*		70*
J0422+32	41.6	380*				≤45*
0620-003	43*	442*	-28	0.311	2.03	90*

THE CATALOG

This paper aims at providing a guide in print to the long and comprehensive full catalog of X-ray binaries at the Web site http://astroa.physics.metu. edu.tr/XRBC/. Below, we present the data and references of HMXBs 0115+634, 0115-737 and LMXBs J0422+32, 0620-003 as examples. The full catalog on the Web will be updated periodically.

The full catalog mainly consists of two parts, hmxb and lmxb, each of which has four files:

(1) The main catalog is composed of hmxb_cat and lmxb_cat. The most recent and/or the most reliable data are given in these files.

(2) All of the cross catalog names of the sources and the names of the optical counterparts are given in hmxb_nam and lmxb_nam.

	Table 6 (sample)				
Name	Names of optical counterparts	Other names of sources			
0115+634	V635 Cas	2E 0115.1+6328			
		H 0115+634, 1H 0115+635			
		3U 0115+63, 4U 0115+63			
		1XRS 01152+634			
		PSR B0115+63.4			
0115-737	Sk 160	SMC X-1			
		2A 0116–737, 3A 0115–737			
		1E 0115.5–7342, 2E 0115.8–7342			
		H 0115–737			
		3U 0115–73, 4U 0115–73			
		1XRS 01157–737			
		SK 160, WW 63			
J0422+32	V518 Per	GRO J0422+32			
		Nova Persei 1992			
0620-003	V616 Mon	A 0620–00			
		N Mon 1917, 1975			

(3) The intensity values (in units of μ Jy, erg cm⁻¹ s⁻², and mCrab), energy bands, dates of observations, names of satellites for these observations, and the references of all these data are given in hmxb_obs and lmxb_obs.

(4) All of the data other than the most recent and/or the most reliable ones are presented in hmxb_rem and lmxb_rem. The values given in intervals are also presented in these two files (e.g., a distance value of 4–5 kpc of a source is given in hmxb_rem or lmxb_rem (not in hmxb_cat or lmxb_cat) even if it is the most recent and/or the most reliable distance value of that source).

The columns in the main catalog (hmxb_cat and lmxb_cat) present: (1) numbers of the sources according to right ascension (RA), (2) names, (3) locations, (4) types (H-black hole, B-burster, P-pulsar, T-transient, Rrecurrent tr., Q-QPO), (5) right ascension RA (2000), (6) declination (dec) (2000), (7) RA (1950), (8) dec (1950), (9) Galactic longitude (l^{II}), (10) Galactic latitude (b^{II}) , (11) orbital period (P_{orb}) , (12) pulse period (P_{pulse}) , (13) remarks on period values (P_{rem}) (shown with asterisks), (14) X-ray temperature kT, (15) spectral index (α) , (16) distance, (17) logarithm of neutral hydrogen column density $(\log N_{\rm H})$, (18) optical absorption (A_V) , (19) X-ray luminosity (L_x) , (20) X-ray to optical luminosity (L_X/L_0) , (21) spectral type (SpecT), (22) V (stellar magnitude), (23) B-V (stellar magnitude), (24) U-B (stellar magnitude), (25) E(B-V) (stellar magnitude), (26) mass of the compact object (M_x) , (27) mass of the

		Table 7 (sample)		
Name	References	Dates	Satellites	Energy bands, keV
0115+634	1994IAUC.59901S	Jan., 1971	Uhuru	26
	1994IAUC.59992W	May 15, 1994	BATSE	20-40
	1997MNRAS.284859N	May 18, 1994	BATSE	20-50
		May 28, 1994	BATSE	20-50
		May 30, 1994	BATSE	20-50
	Intensity, µJy 120	Intensity, erg cm ^{-2} s ^{-1}	Intensity, mCrab	
			45	
			65	
			75	
			180	
0115–737	1977ApJ217543P 60	1971–73	Uhuru	2–6
J0422+32	1995ApJ441786C	Aug. 8, 1992	GRO	20-300
	1992IAUC.55801P		GRO	2-11
	1993A&A280L1S		ROSAT	0.1–2.4
	1995ApJ461351C	1994	BATSE	2–11
			≥3	
		9.3×10^{-9}		
		0	0.3	
		9.3×10^{-9}		
0620-003	1975Nat257656E	Aug. 14, 1975	ArielV	2–18
		Jan. 7, 1976	S3	2-6
		Feb. 1976	ArielV	2-6
	02500	Mar. 1976	ArielV	2-6
	83500		~25	
	300			
	1500			
	0</td <td></td> <td></td> <td></td>			

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Table 8 (sample)

		I I	
Name	References	Other Data	Values of Other Data
0115+634	1978Nat273367C	P _{pulse}	3.61 s
	1979Nat., 282240W	P/P	5.9×10^{-11} s
	1983ApJ 270 711W	Distance	4 4 3 5 4 3 kpc
	1991 ApJ 3751 49N	Distance	5_7 kpc
	1992 ApJ 389 657T	I	$8 \times 10^{36} \text{ erg s}^{-1}$
	1080 A & A 217 115V		$(0.7, 3) \times 10^{37}$ arg s ⁻¹
	1989A&A217113 V		$(0.7-5) \times 10^{-10}$ erg s
	1978ApJ225L71J	M _{opt}	$33 M_{\odot}$
	1981ApJ247222H	Spec. Type	O-Be
		V	$14 \stackrel{m}{.} 5-16 \stackrel{m}{.} 3, 15 \stackrel{m}{.} 5$
		B-V	$1 \frac{m}{2} 4$
0115–737	1986A&A15477T	Energy	2.17 keV
	1984ApJ283546W	$L_{\rm X}/L_{\rm opt}$	1.2-8
	1983ApJ266814M	M _X	$1.6 M_{\odot}$
	1981A&A97134B	M _{opt}	$16 M_{\odot}$
	1995A&A303497V	eont	<0.0007, 0
	1993MNRAS.261337R	K_{ont}	19, 23 km s ^{-1}
	1981A&A.,101184B	K _{ont}	$23-27.5 \text{ km s}^{-1}$
	1984ApJ283249H	$a_1 \sin i$	53.5 light-s
	1982A&A 106 339V		67°
	1983ApJ266814M	E(B-V)	$0^{m}03$
0422+32	1993A&A273L11P	P .	0.05
	1005 A 8 A 207 1020	l orb	0.2137, 0.211, 0.210
	1995A&A297103C	Distance	$\sim 2, >1, 2.4, \leq 2.2 \text{ kpc}$
	1993A&A288L1S	$L_{\rm X}/L_{\rm opt}$	44,~500
	1995PASJ4731K	$L_{\rm X}/L_{\rm opt}$	4-75
	1993A&A276L37C	$L_{\rm X}$	$6.7 \times 10^{30}, 1.5 \times 10^{30}, 8 \times 10^{30}$
	1995ApJ455614F	$\log N_{\rm H}$	21.23
	1995ApJ446L59O	$A_{ m V}$	1 ^m .25
	1996A&A312105C	$M_{\mathbf{Y}}$	$>2.4, 3.57, 4-6, 2.9-6.2 M_{\odot}$
	1995ApJ.,441779V	Mont	$<0.5 M_{\odot}$
	1997ApJ 476L 23V	K	$\sim 340, 300-400 \text{ km s}^{-1}$
	1992PASI 44I 15F	i	$41^{\circ}-43^{\circ} < 59^{\circ} 48^{\circ}$
	1002 ApJ 2001 145P	Space Type	M0 V
	1992ApJ599L145K	Spec. Type	
	1995MINKAS.276L.35C		13–20‴7, 12‴6, 22‴4
	1996MNRAS.282.191D	B-V	$0^{m}_{,3}, 0^{m}_{,38}$
	1995ApJ442786C	E(B-V)	0 3, 0 4
)620–003	1995ApJ442358M	P _{orb}	$0^{d}.323$
	1974A&SS29331A	Distance	0.9–1.1, 0.6–1.4, 0.87 kpc
	1977ApJ212203K	Energy	1.5–3.0 keV
	1986ApJ308110V		10^{38} , 1.6×10^{38} , 6×10^{30} erg s ⁻¹
	1977ApJ 212 209K	$\log N_{\rm H}$	21 59
	1994MNR 4 \$ 268 763 \$	M ₁	$6 \sim 10 \ 27 = 32 \ 319 \ 382 \ M_{\odot}$
	1088 Apl 334 226D		0, -10, 2.7 - 5.2, 5.17 - 5.02 M
		^{IVI} opt	$0.0, 0.50, 0.5-0.8 M_{\odot}$
	1077 A -1 217 1910	v	
	1977ApJ2171810		43 km s^{-1}
	1977ApJ217181O 1983ApJ266L27M	$K_{\rm X}$ $K_{\rm opt}$	43 km s ⁻¹ 457 km s ⁻¹
	1977ApJ217181O 1983ApJ266L27M 1994MNRAS.271L.10S	K _X K _{opt} i	43 km s ⁻¹ 457 km s ⁻¹ 57°
	1977ApJ217181O 1983ApJ266L27M 1994MNRAS.271L.10S 1990ApJ359L47H	K _X K _{opt} <i>i</i> Spec. Type	43 km s ⁻¹ 457 km s ⁻¹ 57° K3, K5V, K4–K5V
	1977ApJ217181O 1983ApJ266L27M 1994MNRAS.271L.10S 1990ApJ359L47H	K_{X} K_{opt} i Spec. Type V	$\begin{array}{c} 43 \text{ km s}^{-1} \\ 457 \text{ km s}^{-1} \\ 57^{\circ} \\ \text{K3, K5V, K4-K5V} \\ 12-18^{m} \end{array}$
	1977ApJ2171810 1983ApJ266L27M 1994MNRAS.271L.10S 1990ApJ359L47H	K_{X} K_{opt} i Spec. Type V $B-V$	43 km s^{-1} 457 km s^{-1} 57° $K3, K5V, K4-K5V$ $12-18^{m}$ $0\stackrel{m}{\cdot}2$

optical companion (M_{opt}) , (28) mass function from Xray observations (f_X) , (29) mass function from optical observations (f_{opt}) , (30) eccentricity of the orbit from optical observations (e_{opt}) , (31) semi-amplitude velocity of the compact object (K_X) , (32) semi-amplitude velocity of the optical companion (K_{opt}) , (33) center-ofmass velocity (γ) , (34) semi-major axis multiplied by inclination $(a_1 \sin i)$, (35) $(a_2 \sin i)$, and (36) inclination (i). An asterisk (*) shows that there are remarks on the related data.

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